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# A theoretical study of the volume dependent Knight shifts of Na and Li

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Abstract. We calculate the Knight shifts, K(V), and spin susceptibilities of the light alkalines Na and Li in the frame of the spin density functional theory (SDA) in the volume range covered by the experiments of Bertani *et al.* All contributions to K have been evaluated rigorously using the data of self-consistent KKR-band-structure calculations. Whereas for Li we get reasonable agreement with experiment, in the case of Na none of the available exchange correlation potentials,  $K_{xx}$ , used in the present work reproduces the minimum of K(V) found in the experiments and we show that moderate changes in  $K_{xx}$  would be able to remove this deficiency.

# 1. Introduction

A great deal of experimental and theoretical work has been devoted to the investigation of the Knight shift, the displacement of the NMR frequency of a nucleus in a condensed system relative to the NMR-frequency of the same nucleus in a diamagnetic reference compound. This frequency shift is caused by the interaction between the magnetic moment of the nucleus and the magnetic field induced by the response of the electrons to an external magnetic field. Following the work of Townes et al (1950) who tried to interpret the experiments solely in terms of valence electron quantities at the Fermi energy, the theory of this phenomenon has been worked out subsequently in more detail: in connection with their work on Li, Cohen et al (1959) and also Gaspari et al (1964) pointed out the significance of the core polarization contribution to the hyperfine Hamiltonian. In systems with non-cubic symmetry there exists a contribution due to the spin-dipole interaction (Carter et al 1977). The importance of the orbital contribution,  $K_{\text{orb}}$  has been recognized by Das and Sondheimer (1960). They showed that in the case of a homogeneous electron gas,  $K_{orb}$  is proportional to the Landau diamagnetism and used this term to explain the Knight shift of Be. While the equivalent expressions for  $K_{\rm orb}$ , due to Stephen (1962) and Hebborn (1962) and derived by considering the orbital shielding of a nucleus in a metal, are extremely hard to evaluate, Ebert et al (1986) gave a tractable prescription for the evaluation of  $K_{orb}$  based upon a formula for the orbital susceptibility,  $\chi_{orb}$ , by Benkowitsch and Winter (1983).

At present, a formulation in the framework of the local spin density functional theory (LSDA) is at hand, expressing all contributions to K in terms of band-structure quantities and potentials derived from LSDA. It is applicable to the whole range of metallic densities.

The results of Ebert *et al* (1986) for some transition metals reveal the importance of  $K_{orb}$  and the tendency of the core polarization contribution,  $K_{cp}$ , to cancel the term  $K_s$ , due to the spin polarization of the valence electrons. The situation is likely to be different in simple metals where s- and p-electrons play the main role and the electronic densities are considerably lower.

The present paper is devoted to a theoretical study of the Knight shift of the light alkalines Na and Li for the wide volume range recently investigated experimentally by Bertani *et al* (1990) and is organized as follows: in section 2 we collect the relevant formulae giving a more detailed account of  $K_{orb}$  since in contrast to the work of Ebert *et al* (1986) we avoid any cluster approximation in the evaluation of this term. Section 3 introduces and characterizes the potentials,  $K_{xc}$ , used to describe enhancement effects and derived from the various exchange–correlation energy functionals available from literature. The main features of the volume dependent band-structures of Na and Li will be presented in section 4 of this work. The results for the volume-dependent Knight shift and the volume-dependent susceptibility, including a detailed discussion, are given in section 5 and we close with a brief summary.

# 2. Formalism

Neglecting spin-orbit coupling, the Knight shift, K, may be written as the sum of a spin contribution,  $K_{spin}$ , and an orbital part,  $K_{orb}$ . We use  $K_{spin}$  in the following form used for example by Ebert *et al* (1986)

$$K_{\rm spin} = (8\pi/3)|m(0)|/|B|. \tag{2.1}$$

Here, m(0) is the spin-magnetization density at the nuclear position caused by the homogeneous external field B. In linear response, the relation between the Fourier component,  $B(q, \omega)$ , of an external magnetic field and the induced spin density,  $m(\rho, \omega)$ , involves the lattice Fourier transform,  $\chi_q$ , of the spin-spin correlation function. It reads

$$\boldsymbol{m}(\boldsymbol{\rho},\omega) = \int \mathrm{d}\boldsymbol{\rho}' \,\boldsymbol{\chi}_{\boldsymbol{q}}(\boldsymbol{\rho},\boldsymbol{\rho}';\omega) \,\mathrm{e}^{\mathrm{i}\boldsymbol{q}\boldsymbol{\rho}'} \,\boldsymbol{B}(\boldsymbol{q},\omega). \tag{2.2}$$

Here, the coordinates  $\rho$ ,  $\rho'$  are restricted to one unit cell. In a SDA-RPA approximation we evaluate  $\chi_q$  by setting up and solving the following integral equation (Stenzel and Winter 1985, 1986)

$$\chi_{q}(\boldsymbol{\rho},\boldsymbol{\rho}';\omega) = \chi_{q}^{p}(\boldsymbol{\rho},\boldsymbol{\rho}';\omega) + \int \mathrm{d}\boldsymbol{\rho}_{1} \,\mathrm{d}\boldsymbol{\rho}_{2} \,\chi_{q}^{p}(\boldsymbol{\rho},\boldsymbol{\rho}_{1};\omega) K_{\mathrm{xc}}(\boldsymbol{\rho}_{1},\boldsymbol{\rho}_{2}) \chi_{q}(\boldsymbol{\rho}_{2},\boldsymbol{\rho}';\omega).$$
(2.3)

While the non-interacting susceptibility,  $\chi_q^p$ , occurring in (2.3) is evaluated using the calculated KKR-band-structure, the kernel,  $K_{\rm xc}$ , is given by the second derivative of the LSDA exchange-correlation energy with repect to the magnetization density (Vosko and Perdew 1975)

$$K_{\mathrm{xc}}(\boldsymbol{\rho},\boldsymbol{\rho}') = \{\delta^2 E_{\mathrm{xc}}[n(\boldsymbol{\bar{\rho}}), |\boldsymbol{m}(\boldsymbol{\bar{\rho}})|]/\delta |\boldsymbol{m}(\boldsymbol{\rho})| |\delta \boldsymbol{m}(\boldsymbol{\rho}')|\}|_{|\boldsymbol{m}|=0}.$$
 (2.4)

Inserting (2.2) into (2.1) we finally obtain for  $K_{soin}$ 

$$K_{\text{spin}} = \frac{8\pi}{3} \lim_{\boldsymbol{q} \to 0} \int d\boldsymbol{\rho}' \, \chi_q(0, \boldsymbol{\rho}'; 0) \, e^{i\boldsymbol{q}\boldsymbol{p}'}.$$
(2.5)

In our practical applications we evaluate only the valence electron contributions,  $K_s$ , to  $K_{spin}$  by using the formalism outlined above. The core electron contribution,  $K_{cp}$ , on the other hand, is obtained with the help of the following formula (Ebert *et al* 1986)

$$K_{\rm cp} = \frac{8\pi}{3} \mu_{\rm B} \frac{1}{|\mathbf{B}|} \left( \sum_{n} |\Psi_{\rm ns\uparrow}(0)|^2 - |\Psi_{\rm ns\downarrow}(0)|^2 \right).$$
(2.6)

The sum in (2.6) extends over all core states whose amplitudes at the nuclei are spindependent due to their interaction with the external field and the spin polarized valence electrons.

To evaluate the orbital contribution, we start with the following relation

$$K_{\rm orb} = \frac{-1}{c\hbar Y_{\rm n} |B|} \int d\rho A_n(\rho) j(\rho)$$
(2.7)

with  $Y_n$  the gyromagnetic ratio,  $A_n = \mu_n \times \rho/\rho^3$  the vector potential of the nuclear magnetic moment and *j* the orbital current induced by *B*. The orbital current density  $j(\rho, \omega)$  created by the Fourier component  $A(q, \omega)$  of an external vector potential reads

$$j_{a}(\boldsymbol{\rho},\omega) = \sum_{\beta} \int d\boldsymbol{\rho}' \, \mathcal{D}_{\boldsymbol{q}}^{\alpha\beta}(\boldsymbol{\rho},\boldsymbol{\rho}';\omega) \, e^{i\boldsymbol{q}\boldsymbol{\rho}'} \, A_{\beta}(\boldsymbol{q},\omega).$$
(2.8)

Inserting (2.8) into (2.7) we obtain

$$K_{\rm orb} = \frac{-1}{c\hbar Y_{\rm n}} \lim_{q\to 0} \sum_{\alpha,\beta} \int d\boldsymbol{\rho} \int d\boldsymbol{\rho}' \, e^{iqp'} \, A_{\rm n\alpha}(\boldsymbol{\rho}) D_{q}^{\alpha\beta}(\boldsymbol{\rho},\boldsymbol{\rho}';0) \, \frac{A_{\beta}(\boldsymbol{q},0)}{|\boldsymbol{B}|}.$$
(2.9)

In our applications we choose B in the y- and q in the z-directions and obtain the following relation

$$K_{\rm orb} = -\frac{1}{c} \int \mathrm{d}\boldsymbol{\rho} \int \mathrm{d}\boldsymbol{\rho}' \frac{z'}{\rho^3} \left( z D_0^{xx}(\boldsymbol{\rho}, \boldsymbol{\rho}'; 0) - x D_0^{xx}(\boldsymbol{\rho}, \boldsymbol{\rho}'; 0) \right) \frac{e^{iqz'}}{q}. \tag{2.10}$$

Instead of employing any cluster approximation as in the work of Ebert *et al* (1986) we express  $D^{\alpha\beta}$  in terms of the lattice transform of the one particle Green's function,  $g_k(\rho, \rho'; \omega)$ , in a similar way as derived for the orbital susceptibility by Götz and Winter (1989). We obtain

$$D_{0}^{d\rho}(\rho, \rho'; \omega) = D_{0}^{d\rho \text{para}} + D^{\alpha \text{dia}}$$

$$= \frac{-2}{\pi} \int \frac{dk}{\Omega_{BZ}} d\varepsilon f(\varepsilon) \lim_{\rho_{1} \to \rho} \lim_{\rho'_{1} \to \rho'} \left[ \left( \frac{\partial}{\partial \rho_{\alpha}} - \frac{\partial}{\partial \rho_{1\alpha}} \right) \right]$$

$$\times \left( \frac{\partial}{\partial \rho'_{\beta}} - \frac{\partial}{\partial \rho'_{1\beta}} \right) \operatorname{Im}(g_{k}(\rho, \rho'; \varepsilon)g_{k}(\rho'_{1}, \rho_{1}; \varepsilon))]$$

$$- \frac{4}{\pi} \operatorname{Im} \int \frac{dk}{\Omega_{BZ}} \int d\varepsilon f(\varepsilon) \operatorname{Im} g_{k}(\rho, \rho'; \varepsilon) \delta(\rho - \rho'). \quad (2.11)$$

In (2.11) only the valence electrons contribute to the paramagnetic part of the correlation function  $D_0^{\alpha\beta}(\rho,\rho;\omega)$  leading to  $K_{orb}^{para val}$ . To  $D^{\alpha dia}$ , however, both the core electrons  $(K_{orb}^{dia core})$  and the valence electrons  $(K_{orb}^{dia val})$  contribute. Assuming that the core charge densities of the reference compound and the metal are the same, only  $K_{orb}^{dia val}$  has to be considered in the calculation of K.

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In our applications we calculate Im  $D_0^{\alpha\beta}(\rho, \rho'; \omega)$  directly whereas Re  $D_0^{\alpha\beta}(\rho, \rho'; 0)$  is obtained by applying the Kramers-Kronig relation. The inclusion of 24 bands guarantees converged results.

# 3. Choice of the exchange correlation potential

In contrast to the bandstructure, the spin part of K turned out to depend sensitively on the approximation to the exchange-correlation energy,  $E_{xc}$ , used, an observation which is hardly surprising since according to (2.3) the second derivative of  $E_{xc}$  with respect to the magnetization density enters the theory of  $K_{spin}$ . In the following we give a short compilation of the different expressions for  $E_{xc}$  which lead to the different  $K_{xc}$  used in the present work. For a more detailed discussion we refer the reader to the original papers.

The local functionals are of the form

$$E_{\rm xc}[n,m] = \int dr \, n(r) \varepsilon_{\rm xc}(n(r),m(r)) \tag{3.1}$$

with  $\varepsilon_{xc} = \varepsilon_x(n, \zeta) + \varepsilon_c(n, \zeta)$  and  $\zeta = |m|/n$  the relative magnetization. They differ in the approximation to the correlation part  $\varepsilon_c$ .

Von Barth and Hedin (1972) (vBH) calculated  $\varepsilon_c(r_s, \zeta)$  for a number of densities  $(n^{-1} = 4\pi/3r_s^3)$  in the metallic range and some values of  $\zeta$ . Their applications are based on a RPA-type of approximation for the dielectric function. They gave the following parametrization of their results

$$\varepsilon_{\rm c}(r_{\rm s},\zeta) = \varepsilon_{\rm c}^{\rm vBH}(r_{\rm s},0) + [\varepsilon_{\rm c}^{\rm vBH}(r_{\rm s},1) - \varepsilon_{\rm c}^{\rm vBH}(r_{\rm s},0)]f(\zeta)$$
(3.2)

with  $\varepsilon_c^{\text{vBH}}(r_s, 1/0) = -c^{\text{F/P}}G(r_s/r^{\text{F/P}})$  and  $G(x) = (1 + x^3) \ln(1 + 1/x) - x^2 + x/2 - \frac{1}{3}$ . Both Janak *et al* (1975) (JMW) and Gunnarsson and Lundqvist (1976) (GL) and also Rajagopal (1980) (R) who go beyond the RPA end up with the same functional form, but their parameters  $c^{\text{F/P}}$  and  $r^{\text{F/P}}$  are somewhat different.

We also used the expressions given by Perdew and Zunger (1981) (PZ) and Vosko *et al* (1980) (VWN). They are based on the Monte Carlo results of Ceperley and Alder (1980) (CA) for  $\varepsilon_c(r_s, 1)$  and  $\varepsilon_c(r_s, 0)$ , respectively, at some values for  $r_s$ . PZ and VWN derive interpolation formulae for these quantities approaching the correct high density limit as well. To describe the situation of partial magnetization, PZ use the following approximate relation analogous to (3.2).

$$\varepsilon_{\rm c}^{\rm PZ}(r_{\rm s},\zeta) = \varepsilon_{\rm c}^{\rm CA}(r_{\rm s},0) + [\varepsilon_{\rm c}^{\rm CA}(r_{\rm s},1) - \varepsilon_{\rm c}^{\rm CA}(r_{\rm s},0)]f(\zeta). \tag{3.3}$$

The derivation of  $\varepsilon_c^{VWN}(r_s, \zeta)$  is more delicate, but both  $\varepsilon_c^{VWN}(r_s, \zeta)$  and  $\varepsilon_c^{PZ}(r_s, \zeta)$  are somewhat uncertain due to the lack of Monte-Carlo results for partial magnetizations.

In our application of a non-local potential,  $K_{xc}^{nl}$ , we followed the lines of Geldart and Taylor (1970a, b) (GT) approximating  $K_{xc}(\rho, \rho')$  by its value  $K_{xc}^{h}(\bar{n}, |\rho - \rho'|)$  for a homogeneous electron gas of density  $\bar{n} = \frac{1}{2}(n(\rho) + n(\rho'))$  whose Fourier transform,  $K_{xc}^{h}(\bar{n}, |q|)$  and  $\chi^{P}(\bar{n}, |q|)$  are related to the interacting and non-interacting susceptibilities  $\chi(\bar{n}, |q|)$  and  $\chi^{P}(\bar{n}, |q|)$ , respectively by the following relation

$$K_{xc}^{h}(n, |q|) = \chi^{P}(n, |q|)^{-1} - \chi(n, |q|)^{-1}.$$
(3.4)

We approximated  $\chi^{P}$  and  $\chi$  by the same diagram as Geldart and Taylor. For comparison we also probed the local limit by replacing  $K_{xc}^{GT}(\bar{n}, |q|)$  with  $K_{xc}^{GT}(\bar{n}, 0)$ .

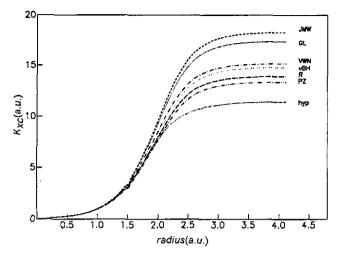


Figure 1. The radial dependence of some of the  $K_{x,s}$  presented in section 3 for Na at zero pressure. The potential  $K_{x,s}^{hyp}$  will be introduced in section 5.

In a series of papers, an extension of the LSDA has been worked out (Langreth and Perdew 1977, 1980) which is based on a modified gradient expansion removing the shortcomings of the gradient expansion originally proposed by Hohenberg and Kohn (1964). In various applications this formulation leads to major improvements in the description of non-magnetic as well as magnetic groundstate properties (Bagno *et al* 1989). We worked with the two potentials  $K_{xc}^{LMH}$  and  $K_{xc}^{PW}$  derived according to (2.4) from the expressions for  $E_{xc}$  given by Langreth and Mehl (1983) and Hu and Langreth (1985) and Perdew (1986a, b) and Perdew and Wang (1986), respectively.

For illustration we show the radial dependencies of the potentials,  $K_{xc}$ , within the Wigner-Seitz sphere resulting for some of the approximations described above, for sodium in figure 1 and for lithium in figure 2. These figures show that the potentials

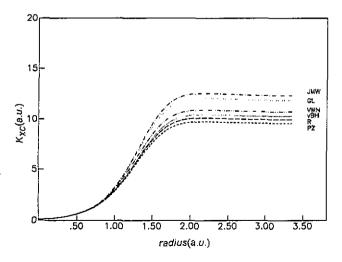


Figure 2. The radial dependence of some of the  $K_{xc}$ s presented in section 3 for Li at zero pressure.

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obtained with the various approaches differ substantially, especially near the ws-sphere boundary. In the case of sodium which has the lower electronic density these variations are more pronounced than in the case of lithium.

Whereas all the other potentials increase monotonically as a function of the local values of  $r_s$ , the  $K_{xc}s$  derived from the modified gradient expansions show a pronounced maximum in the range where the core charge density falls to zero and the valence electron density starts to dominate. This feature is caused by the dependence of this kind of potential,  $K_{xc}$ , on both the first and the second derivative of the charge density.

#### 4. Features of the volume dependent band structure

We performed self-consistent KKR-band-structure calculations for Na and Li at some lattice constants (column 1 of table 1(2) for Na(Li)) in the pressure range covered by the experiments of Bertani *et al* (1990), and yielding the energies and state-vectors of the 24 lowest valence bands. In all cases concerning Na the occupied part of the lowest bands is of parabolic shape. The DOS,  $n(\varepsilon)$ , is dominated by  $s(n_0)$ - and  $p(n_1)$ -partial DOS contributions. Also, a non-negligible *d*-admixture  $(n_2)$  exists. Tables 1 and 2 show these quantities at the Fermi energy  $(\varepsilon_F)$ . In the case of Na the occupied part of the valence band, and also  $\varepsilon_F$  and  $n(\varepsilon_F)$  show jellium-like behaviour.  $\varepsilon_F$  increases, whereas  $n(\varepsilon_F)$ decreases monotonically with decreasing volume. These relations are—also quantitatively—reasonably well described by the formulae of the jellium model

$$\varepsilon_{\rm F} = (3\pi^2)^{2/3} \Omega_{\rm WS}^{-2/3} \tag{4.1}$$

$$n(\varepsilon_{\rm F}) = [(3\pi^2)^{1/3}/2\pi^2]\Omega_{\rm WS}^{2/3} \tag{4.2}$$

with  $\Omega_{WS}$  the volume of the ws-cell.

However, as can be already deduced from the partial DOS, the electron-state vectors deviate markedly from plane waves. This observation applies to both the Bloch coefficients and the radial parts,  $R_1(\rho; \varepsilon)$  of the single-site wave functions. The value of  $R_0^2(0; \varepsilon_F)$  which turns out to be crucial for the magnitude of  $K_s$  is given in column 9 of tables 1 and 2. It rises monotonically with decreasing volume.

In the case of Li the occupied part deviates significantly from a jellium-like behaviour in all respects: the differences to a parabolic band shape and also the anisotropies are non-negligible,  $n(\varepsilon_{\rm F})$  depends non-monotonically on the volume and the features at  $\varepsilon_{\rm F}$ are structure dependent (volumes 1 to 3 of table 2 refer to the BCC-structure, whereas row 4 refers to the FCC-structure). As can be seen from column 9 of table 2  $R_{\rm s}^2(0; \varepsilon_{\rm F})$ increases slightly with pressure in the BCC-phase.

The higher bands and state vectors of these substances which are needed for the calculation of  $K_{orb}$  (2.11) do not show any resemblance to plane waves and cannot be approximated by resorting to any model.

а	$V/V_0$	€ <sub>F</sub>	$n(\varepsilon_{\rm F})$	$n_0(\varepsilon_F)$	$n_1(\varepsilon_F)$	$n_2(\varepsilon_{\rm F})$	$n_3(\varepsilon_{\rm F})$	$R_0^2(0;\varepsilon_{\Gamma})$
8.126	1.00	0.2164	6.569	2.630	3.283	0.608	0.048	18.228
7.8456	0.90	0.2451	6.196	2.386	3.176	0.586	0.048	19.292
7,600	0.82	0,2716	5.866	2.203	3.052	0.566	0.045	20.962
7.0391	0.65	0.3499	5.176	1.825	2.782	0.327	0.042	24.920

Table 1. Volume dependent band-structure quantities for Na.

a	$V/V_0$	ε <sub>F</sub>	$n(\varepsilon_{\rm F})$	$n_0(\varepsilon_{\rm F})$	$n_1(\varepsilon_{\rm F})$	$n_2(\varepsilon_{\rm F})$	$n_3(\varepsilon_F)$	$R_0^2(0;\varepsilon_{\rm F})$
6.6338	1.00	0.2961	6.965	1.504	4.997	0.401	0.063	6.243
6.4984	0.94	0.3152	6.402	1.522	4.471	0.357	0.052	6.547
6.1255	0.79	0.3641	6.948	1.259	5.255	0.361	0.073	7.215
7.4964	0.72	0.3719	6.144	1.199	4.563	0.329	0.053	8.356

Table 2. Volume dependent band structure quantities for Li.

#### 5. Results and discussion

#### 5.1. Na

In table 3 we collect our results for K(V) and  $K_{rel}(V) = K(V)/K(V_0)$ . These quantities are indexed according to the potential,  $K_{\rm xc}$ , used in their computation. The values of the individual terms defined in section 2 are also given. It is obvious that the spin contribution,  $K_s$ , of the valence electron dominates  $K_{cp}$  and  $K_{orb}$  in the whole volume range. This is the case for both K and  $K_{rel}$ . As already suggested by figure 1, table 3 confirms the strong dependence of the results on  $K_{\rm xc}$ . As a common feature we observe a monotonic decrease of  $K_{rel}$  with decreasing volume, smoothening at higher pressures. Its magnitude, however, varies considerably between the highest value provided by  $K_{xc}^{PZ}$  and  $K_{xc}^{OT}$  on the one side and the lowest value which is due to  $K_{xc}^{JMW}$ . The differences between the absolute values,  $K(V_0)$ , are still more pronounced:  $K^{\text{LMH}}(V_0)$ , for example, is about 30% above  $K^{PZ}(V_0)$  and  $K^{GT}(V_0)$ . All the calculated values of  $K_{rel}(V)$  fall below the experimental results of Bertani et al (1990) and fail to reproduce the minimum at V/ $V_0 = 0.84$  found by these authors.  $K_{xc}^{PZ}$  and  $K_{xc}^{GT}$  lead to the best results with respect to both  $K(V_0)$  and  $K_{rel}(V)$ , whereby  $K(V_0)$  is about 15% above the experimental number. We extended our calculations to the spin susceptibility,  $\chi$ , and the numbers for  $\chi_{rel}(V) =$  $\chi(V)/\chi(V_0)$  are displayed in column 9 of table 3. Whilst all the potentials,  $K_{\rm xc}$ , yield a steep increase of different magnitude at high pressures, the slopes of  $\chi_{rel}$  at small volume changes vary even in sign. Our low pressure values (e.g.  $\chi_{rel}^{PZ}(0.9) = 1.01$ ) lie below the experimental data of Kushida et al (1976) ( $\chi_{rel}(0.9) = 1.033$ ) while no experimental investigations of  $\chi$  in the high pressure regime are known to us.

The results of table 3 may be compared to other theoretical work. Manninen and Jena (1980) calculated  $K_s(V)$  in the range of small pressures for a Na ion immersed into a Fermi sea of a homogeneous electron gas. The deviations of their  $K_s(V)$  from experiment are similar to ours. Wilk and Vosko (1981) take account of the band-structure by evaluating the effect of an external magnetic field on the bands and the wave-vectors in the first order. In accordance with us they find that the spin contribution,  $K_s$ , dominates and also their  $K(V)/K(V_0)$  is close to our value derived from  $K_{xc}^{VWN}$ . On the other hand, their  $K(V_0)$  is near to experiment. As these authors give no details on their band-structure data, a more precise comparison is impossible.

Our results for  $K_s$  may be made more transparent with the help of the following simplified formula (Moruzzi *et al* 1978)

$$K_{\rm s} = (8\pi/3)Sn_{\rm s}(\varepsilon_{\rm F})R_{\rm s}^2(0;\varepsilon_{\rm F}) \tag{5.1}$$

with S the Stonerenhancement,  $n_s(\varepsilon_F)$  the partial s-DOS and  $R_s(0; \varepsilon_F)$  the radial part of the single-site s-wave function at the nucleus. It reproduces the numbers of table 3 roughly. Whereas,  $R_s^2$  increases almost linearly with pressure, the factor  $Sn_s$  decreases

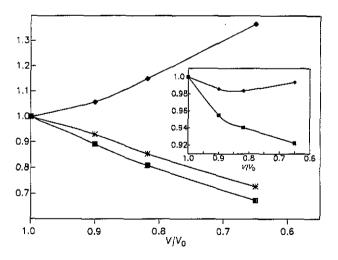
**Table 3.** Results for the different contributions to the Knight shift of Na as a function of reduced volume for the potentials presented in section 3 (all values in %).  $K_{xc}^{hyp}$  will be introduced in section 5. Also the quantities  $K_{rel}(V) = K(V)/K(V_0)$  and  $\chi_s(V)/\chi_s(V_0)$  are listed.

Potential	а	$V/V_0$	Κ	$K_{\rm rel}$	Ks	K <sub>cp</sub>	$K_{ m orb}^{ m dia val}$	K para vai orb	X ret
GL	8.126	1.00	0.1495	1.000	0.1528	-0.0032	$-0.69 \times 10^{-3}$	$0.57 \times 10^{-3}$	1,000
	7.8456	0,90	0.1418	0.948	0.1451	-0.0032	$-0.71 \times 10^{-3}$	$0.60 \times 10^{-3}$	0.999
	7.600	0.82	0.1394	0.932	0.1428	-0.0033	$-0.74 \times 10^{-3}$	$0.63  imes 10^{-3}$	1.052
	7.0391	0.65	0.1343	0.898	0.1379	-0.0035	$-0.81 \times 10^{-3}$	$0.71 \times 10^{-3}$	1.124
GT	8.126	1.00	0.1327	1.000	0.1345	-0.0017	$-0.69 \times 10^{-3}$	$0.57 \times 10^{-3}$	1.000
(local)	7.8456	0.90	0.1274	0.960	0.1293	-0.0018	$-0.71 \times 10^{-3}$	$0.60 \times 10^{-3}$	1.013
	7.600	0.82	0.1261	0.950	0.1281	-0.0019	$-0.74 \times 10^{-3}$	$0.63 \times 10^{-3}$	1.073
	7.0391	0.65	0.1236	0.931	0.1259	-0.0022	$-0.81 \times 10^{-3}$	$0.71 \times 10^{-3}$	1.170
GT	8.126	1.00	0.1311	1.000	0.1329	-0.0017	$-0.69 \times 10^{-3}$	$0.57 \times 10^{-3}$	1.000
(non-local)	7.8456	0.90	0.1257	0.959	0.1276	-0.0018	$-0.71 \times 10^{-3}$	$0.60 \times 10^{-3}$	1.011
	7.600	0.82	0,1246	0.950	0.1266	-0.0019	$-0.74 \times 10^{-3}$	$0.63 \times 10^{-3}$	1.068
	7.0391	0.65	0.1221	0.931	0.1243	-0.0021	$-0.81 \times 10^{-3}$	$0.71 \times 10^{-3}$	1.154
JMW	8.126	1.00	0.1541	1.000	0.1578	-0.0036	$-0.69 \times 10^{-3}$	$0.57 \times 10^{-3}$	1.000
	7.8456	0.90	0.1458	0.946	0.1494	-0.0035	$-0.71 \times 10^{-3}$	$0.60 \times 10^{-3}$	0.995
	7.600	0.82	0.1429	0.927	0.1467	-0.0037	-0.74 × 10 <sup>-3</sup>	$0.63 \times 10^{-3}$	1,046
	7.0391	0.65	0.1371	0.890	0.1410	-0.0038	$-0.81 \times 10^{-3}$	$0.71 \times 10^{-3}$	1.111
LMH	8.126	1.00	0.1813	1.000	0.1841	-0.0027	$-0.69 \times 10^{-3}$	$0.57 \times 10^{-3}$	1.000
	7.8456	0.90	0.1731	0.955	0.1760	-0.0028	$-0.71 \times 10^{-3}$	$0.60 \times 10^{-3}$	1.009
	7.600	0.82	0.1686	0.930	0.1716	-0.0029	$-0.74 \times 10^{-3}$	$0.63 \times 10^{-3}$	1.060
	7.0391	0.65	0.1649	0.910	0.1682	-0.0032	$-0.81 \times 10^{-3}$	$0.71 \times 10^{-3}$	1.151
PZ	8.126	1.00	0.1312	1.000	0.1330	-0.0017	$-0.69 \times 10^{-3}$	$0.57 \times 10^{-3}$	1.000
	7.8456	0.90	0.1261	0.961	0.1280	-0.0018	$-0.71 \times 10^{-3}$	$0.60 \times 10^{-3}$	1.011
	7.600	0.82	0.1246	0.950	0.1266	-0.0019	$-0.74 \times 10^{-3}$	$0.63 \times 10^{-3}$	1.073
	7.0391	0.65	0.1221	0.931	0.1244	-0.0022	$-0.81 \times 10^{-3}$	$0.71 \times 10^{-3}$	1.169
R	8.126	1.00	0.1335	1.000	0.1355	-0.0019	$-0.69 \times 10^{-3}$	$0.57 \times 10^{-3}$	1.000
	7.8456	0.90	0.1280	0.959	0.1301	-0.0020	$-0.71 \times 10^{-3}$	$0.60 \times 10^{-3}$	1.012
	7.600	0.82	0.1265	0.948	0.1287	-0.0021	$-0.74 \times 10^{-3}$	$0.63 \times 10^{-3}$	1.071
	7.0391	0.65	0.1238	0.927	0.1263	-0.0024	$-0.81 \times 10^{-3}$	$0.71 \times 10^{-3}$	1.165
vBH	8.126	1.00	0.1370	1.000	0.1392	-0.0021	$-0.69 \times 10^{-3}$	$0.57 \times 10^{-3}$	1.000
	7.8456	0.90	0.1308	0.955	0.1331	-0.0022	$-0.71 \times 10^{-3}$	$0.60 \times 10^{-3}$	1.007
	7.600	0.82	0.1290	0.942	0.1314	-0.0023	$-0.74 \times 10^{-3}$	$0.63 \times 10^{-3}$	1.064
	7.0391	0.65	0.1256	0.917	0.1282	-0.0025	$-0.81 \times 10^{-3}$	$0.71 \times 10^{-3}$	1,124
VWN	8.126	1.00	0.1390	1.000	0.1412	-0.0021	$-0.69 \times 10^{-3}$	$0.57 \times 10^{-3}$	1.000
	7.8456	0.90	0.1330	0.957	0.1353	-0.0022	$-0.71 \times 10^{-3}$	$0.60 \times 10^{-3}$	1.009
	7.600	0.82	0.1314	0.945	0.1338	-0.0023	$-0.74 \times 10^{-3}$	$0.63 \times 10^{-3}$	1.067
	7.0391	0.65	0.1282	0.922	0.1308	-0.0025	$-0.81 \times 10^{-3}$	$0.71 \times 10^{-3}$	1.155
hyp	8.126	1.00	0.1211	1.000	0.1228	-0.0016	$-0.69 \times 10^{-3}$	$0.57 \times 10^{-3}$	1.000
	7.8456	0.90	0.1194	0.986	0.1212	-0.0017	$-0.71 \times 10^{-3}$	$0.60 \times 10^{-3}$	1.039
	7.600	0.82	0.1192	0.984	0.1211	-0.0018	$-0.74 \times 10^{-3}$	$0.63 \times 10^{-3}$	1.111
	7.0391	0.65	0.1212	1.001	0.1235	-0.0022	$-0.81 \times 10^{-3}$	$0.71 \times 10^{-3}$	1.257

fast enough to cause the negative slope of  $K_s$  in the whole volume range. The high sensitivity of  $K_s$  upon changes in the potential  $K_{yr}$ , is due to the occurrence of the factor S in (5.1). Because the number for  $K_s(V_0)$ , (0.1384), obtained by Moruzzi et al (1978) using the lattice constant a = 7.7 au is comparable to ours (referring to a = 7.6 au) and the APW-bandstructure calculations of Papaconstantopoulos (1986) for zero pressure agree with our KKR-results with respect to  $\varepsilon_{\rm F}$  and  $n(\varepsilon_{\rm F})$  while their  $n_{\rm s}(\varepsilon_{\rm F})$  is somewhat higher, our band-structure data cannot be blamed for the theoretical overestimation of the experimental value of  $K_s(V_0)$ . Also, relativistic effects (Tterlikkis *et al* 1969) which turned out to be important for the heavier alkalines (changes of 50% in K, for Cs) and non-negligible in the case of Rb ( $\sim 5\%$ ) should be unimportant for Na. In view of the large differences between the results obtained with the various available potentials,  $K_{xc}$ , and the discrepancies to the experiment it may be worthwhile investigating the changes of  $K_{xc}$  necessary to get better agreement. In this connection it is important to remark that both S and the matrix elements of  $K_{xc}$ , entering the exact RPA-LSDA formalism are determined by the weakly space-dependent part of  $K_{\rm xc}(\rho)$  in the outer region of the wssphere, a feature that is due to the dominance of s- and p-type partial waves in the Bloch states. This is in variance with the situation in transition metals where the average electron density is higher and the most important partial waves are of d-character peaking at smaller radii. We introduced the hypothetical potential,  $K_{xc}^{hyp}$ , whose values lie somewhat below those of  $K_{xc}^{PZ}$ . In the flat part near the ws-sphere boundary and within the pressure range considered  $(0.65 \le V/V_0 \le 1)$  the relation between  $K_{re}^{hyp}$  and  $K_{\rm xc}^{\rm PZ}$  is given by the following formula

$$K_{\rm xc}^{\rm hyp}(\rho, (V/V_0)) = K_{\rm xc}^{\rm PZ}(\rho, (V/V_0))(1 - 0.47((V/V_0) - 0.65)).$$
(5.2)

So  $K_{xc}^{hyp}$  coincides with  $K_{xc}^{PZ}$  at the upper end of the pressure range, where the average electron density comes near to that of Li, while the differences are largest at zero pressure.  $K_{xc}^{hyp}$  is also drawn in figure 1. Its deviation from  $K_{xc}^{PZ}$  is smaller than the scatter among the other potentials. Our results based on  $K_{xc}^{hyp}$  are shown in table 3.  $K_{rep}^{hyp}(V)$ 



**Figure 3.** The values of the quantities  $X(V) = R_s^2(0, \varepsilon_F; V)/R_s^2(0, \varepsilon_F; V_0)$ :  $\blacklozenge$  and  $Z_1 = S(V)n_s(\varepsilon_F; V)/(S(V_0)n_s(\varepsilon_F; V_0))$ :  $\blacksquare$  for Na resulting from  $K_{xc}^{PZ}$ . The curve marked \* shows the latter quantity  $(Z_2(V))$  as obtained from the hypothetical potential  $K_{xc}^{hyp}$ . Insert:  $K_{rel}(V) = X(V)Z_1(V)$ :  $\blacksquare K_{rel} = X(V)Z_2(V)$ :  $\blacklozenge$ .

agrees with the experiments of Bertani *et al* (1990) and also the absolute values,  $K^{hyp}(V)$ , show improvement: their overall deviation from experiment has diminished to 7%. Figure 3 illustrates the effect of  $K_{xc}^{hyp}$  in terms of the simplified relation according to (5.1): As a consequence of the smoother volume dependence of S the negative slope of  $n_sS$  has slightly decreased, leading to an even qualitative change in the behaviour of  $K_{rel}(V)$ . The use of  $K_{xc}^{hyp}$  also yields improvement on the susceptibility data, as a comparison of  $\chi_{rel}^{hyp}$  (table 3) with the low pressure experiments of Kushida *et al* (1976) shows. Summarizing, we stress that—in contrast to the other potentials— $K_{xc}^{hyp}$  is not really founded from physical grounds. Its purpose is only to illustrate the high sensitivity of K(V) on  $K_{xc}$  and to give some idea of the changes to  $K_{xc}$  necessary to bring the theoretical results nearer to experiment. Since both  $K_{xc}^{PZ}$  and  $K_{xc}^{VWN}$  date back to the Monte-Carlo calculations of Ceperley and Alder (1980) performed for the fully spin-polarized and the paramagnetic case only, computations of such a kind for the partly spin-polarized situation would be highly desirable in order to remove possible uncertainties in the corresponding potentials  $K_{xc}$ .

# 5.2. Li

 $K(V_0)$ ,  $K_{rel}(V)$  and  $\chi_{rel}(V)$  have been evaluated for four lattice constants, the last one referring to the FCC-phase (table 4). In accordance with figure 2 the results show less dependence on  $K_{xc}$  than in the case of Na. The maximum difference for  $K(V_0)$  occurs between  $K^{PZ}(V_0)$  and  $K^{JMW}(V_0)$  amounting to 21%. Common to all potentials is a nonlinear increase of  $K_{rel}(V)$  in the higher pressure regime to numbers around 1.05 and a jump at the transition to the FCC-phase to values near 1. In contrast to Na, besides  $K_s$ ,  $K_{cp}$  contributes non-negligibly to K. It has a negative sign and its magnitude is about  $0.2K_s$ . Its volume dependence is more pronounced than that of  $K_s$ , leading to a significant influence of  $K_{cp}$  on  $K_{rel}(V)$ .  $K_{orb}$  on the other hand turns out to be quite small, whereby  $K_{orb}^{daval}$  which has negative sign tends to cancel  $K_{orb}^{paraval}$ .

In terms of the simplified formula (5.1), which is a poorer approximation to K(V) in the case of Li than for Na, the pronounced increase of  $K_{rel}(V)$  is due to the rise of both  $R_s^2(0; \varepsilon_F)$  and S. These factors would lead to  $K_{rel}(0.79) \approx 1.1$ , a value which is considerably reduced by the volume dependence of  $K_{cp}$ . The jump of  $K_{rel}$  at the phase transition on the other hand is caused by the smaller values of  $n_s$  and S in the FCC-phase, whereas  $R_s^2(0; \varepsilon_F)$  is slightly above the number in the BCC-phase. Our results may be compared with calculations based on the model of an ion immersed into a homogeneous electron gas. Zaremba and Zobin (1980) use  $K_{xc}^{GL}$  to evaluate  $K(V_0)$ . Their data for  $K_s$  and  $K_{cp}$  deviate from ours by 12% and 35%, respectively, whereas—due to compensating effects— $K(V_0)$  deviates only by 6%. The calculations of Manninen and Jena (1980) extending to reduced volumes  $V/V_0 = 0.9$  yield a steep decrease of  $K_{rel}(K_{rel}(0.9) = 0.95)$  stressing the importance of band-structure effects in this system. Wilk and Vosko (1981) treated volume changes up to  $V/V_0 = 0.9$ . Their  $K_{rel}(0.9)$  using  $K_{xc}^{VWN}$  is near to ours, but the initial slope of  $K_{rel}$  at  $V/V_0 \sim 1$  is slightly negative and  $K(V_0)$  falls 9% below our value.

The experiments of Kushida and Murphy (1980) (performed for  $V/V_0$  down to ~0.9) agree with those of Bertani *et al* (1990) with respect to  $K(V_0)$ , but yield a somewhat smaller initial slope for  $K_{rei}(V)$ .

Our calculated  $K(V_0)$  is near to experiment,  $K_{rel}(V)$  follows the curve of Kushida and Murphy (1980) at low pressures and for all potentials,  $K_{xc}$ , used,  $K_{rel}(V)$  is compatible with the results of Bertani *et al* in the range of higher pressures and within the BCC-phase.

**Table 4.** Results for the different contributions to the Knight shift of Li as a function of reduced volume for different potentials,  $K_{xx}$ .

Potential	а	$V/V_0$	K	K <sub>rei</sub>	 K,	K <sub>cp</sub>	K diaval	K erb	$\chi_{rei}$
GL.	6.6338 6.4984 6.1255	1.00 0.94 0.79	0.028 79 0.028 38 0.030 20	1.000 0.986 1.049	0.039 08 0.038 02 0.04476	-0.010 04 -0.009 45 -0.014 31	$-0.84 \times 10^{-3} \\ -0.81 \times 10^{-3} \\ -0.92 \times 10^{-3}$	$\begin{array}{c} 0.59 \times 10^{-3} \\ 0.63 \times 10^{-3} \\ 0.67 \times 10^{-3} \end{array}$	1.000 1.092 1.127
	7.4964	0.72	0.028 94	1.005	0.041 60	-0.012 29	$-0.90 \times 10^{-3}$	$0.54 \times 10^{-3}$	1.027
Gт (local)	6.6338 6.4984 6.1255 7.4964	1.000 0.94 0.79 0.72	0.024 76 0.024 80 0.025 84 0.024 83	1.000 1.002 1.044 1.003	0.031 52 0.031 14 0.035 41 0.033 83	-0.006 50 -0.006 15 -0.009 32 -0.008 64	$\begin{array}{c} -0.84 \times 10^{-3} \\ -0.81 \times 10^{-3} \\ -0.92 \times 10^{-3} \\ -0.90 \times 10^{-3} \end{array}$	$\begin{array}{c} 0.59 \times 10^{-3} \\ 0.63 \times 10^{-3} \\ 0.67 \times 10^{-3} \\ 0.54 \times 10^{-3} \end{array}$	1.000 1.076 1.106 1.035
GT (non-local)	6.6338 6.4984 6.1255 7.4964	1.00 0.94 0.79 0.72	0.023 83 0.023 88 0.024 91 0.023 86	1.000 1.002 1.043 1.001	0.030 34 0.029 99 0.033 92 0.032 23	-0.006 26 -0.005 93 -0.008 76 -0.008 01	$\begin{array}{c} -0.84\times10^{-3}\\ -0.81\times10^{-3}\\ -0.92\times10^{-3}\\ -0.90\times10^{-3} \end{array}$	$\begin{array}{c} 0.59 \times 10^{-3} \\ 0.63 \times 10^{-3} \\ 0.67 \times 10^{-3} \\ 0.54 \times 10^{-3} \end{array}$	1.000 1.056 1.080 0.996
MM	6.6338 6.4984 6.1255 7.4964	1.00 0.94 0.79 0.72	0.029 91 0.029 35 0.031 46 0.030 13	1.000 0.981 1.052 1.007	0.041 14 0.039 82 0.047 39 0.043 62	-0.010 97 -0.010 28 -0.015 67 -0.013 13	$\begin{array}{c} -0.84\times10^{-3} \\ -0.81\times10^{-3} \\ -0.92\times10^{-3} \\ -0.90\times10^{-3} \end{array}$	$\begin{array}{c} 0.59 \times 10^{-3} \\ 0.63 \times 10^{-3} \\ 0.67 \times 10^{-3} \\ 0.54 \times 10^{-3} \end{array}$	1.000 1.096 1.134 1.023
P2	6.6338 6.4984 6.1255 7.4964	1.00 0.94 0.79 0.72	0.024 56 0.024 67 0.025 47 0.024 68	1.000 1.004 1.037 1.005	0.031 26 0.031 00 0.034 92 0.033 67	-0.006 45 -0.006 14 -0.009 19 -0.008 62	$\begin{array}{c} -0.84 \times 10^{-3} \\ -0.81 \times 10^{-3} \\ -0.92 \times 10^{-3} \\ -0.90 \times 10^{-3} \end{array}$	$\begin{array}{c} 0.59 \times 10^{-3} \\ 0.63 \times 10^{-3} \\ 0.67 \times 10^{-3} \\ 0.54 \times 10^{-3} \end{array}$	1.000 1.072 1.110 1.038
R	6.6338 6.4984 6.1255 7.4964	1.00 0.94 0.79 0.72	0.025 12 0.025 17 0.026 14 0.025 20	1.000 1.002 1.041 1.003	0.032 28 0.031 94 0.036 26 0.034 80	-0.006 91 -0.006 58 -0.009 87 -0.009 23	$-0.84 \times 10^{-3}$ $-0.81 \times 10^{-3}$ $-0.92 \times 10^{-3}$ $-0.90 \times 10^{-3}$	$\begin{array}{c} 0.59 \times 10^{-3} \\ 0.63 \times 10^{-3} \\ 0.67 \times 10^{-3} \\ 0.54 \times 10^{-3} \end{array}$	1.000 1.076 1.106 1.039
vBH	6.6338 6.4984 6.1255 7.4964	1.00 0.94 0.79 0.72	0.025 72 0.025 70 0.026 67 0.025 99	1.000 0.999 1.037 1.010	0.033 26 0.032 80 0.037 21 0.035 59	-0.007 29 -0.006 92 -0.010 29 -0.009 24	$\begin{array}{c} -0.84 \times 10^{-3} \\ -0.81 \times 10^{-3} \\ -0.92 \times 10^{-3} \\ -0.90 \times 10^{-3} \end{array}$	$\begin{array}{c} 0.59 \times 10^{-3} \\ 0.63 \times 10^{-3} \\ 0.67 \times 10^{-3} \\ 0.54 \times 10^{-3} \end{array}$	1.000 1.073 1.102 1.030
VWN	6.6338 6.4984 6.1255 7.4964	1.00 0.94 0.79 0.72	0.026 88 0.026 84 0.028 33 0.026 99	1.000 0.999 1.054 1.004	0.034 75 0.034 24 0.039 51 0.037 45	-0.007 62 -0.007 21 -0.010 93 -0.010 09	$\begin{array}{c} -0.84 \times 10^{-3} \\ -0.81 \times 10^{-3} \\ -0.92 \times 10^{-3} \\ -0.90 \times 10^{-3} \end{array}$	$\begin{array}{c} 0.59 \times 10^{-3} \\ 0.63 \times 10^{-3} \\ 0.67 \times 10^{-3} \\ 0.54 \times 10^{-3} \end{array}$	1.000 1.086 1.119 1.040

Also, the jump to lower values of  $K_{rel}(V)$  at the phase transition is in qualitative agreement with experiment, but its magnitude is below the data of Bertani *et al.* As a comparison of  $K^{GT}(V_0)$  to  $K^{GTnl}(V_0)$  shows, non-localities in  $K_{xc}$  give rise to changes in the percentage range.

The behaviour of the spin susceptibility is characterized by a steep rise of  $\chi_{rel}$  at high pressures. Its slope at small compressions lies considerably above the values obtained for  $K_{rel}$  and is—especially in the case of  $\chi_{rel}^{VWN}$  and  $\chi_{rel}^{R}$ —near to the low pressure experiments of Kushida *et al* (1976) who obtain  $\chi_{rel}(0.94) = 1.06$ . At the phase transition, the theoretical  $\chi_{rel}$  jumps from 1.4 to 1.25, an effect which is due to the lower values of  $n(\varepsilon_F)$  and S in the FCC-phase. The corresponding results of Wilk *et al* (1979) who calculated  $\chi$  up to  $V/V_0 = 0.9$  using the potentials of Shastry as well as of Keiser and Wu and employing the variational principle to include enhancement effects show the same tendency as ours.

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Summarizing, one can say that the general agreement between theory and experiment for Li is better than for Na. In this connection it might be interesting to keep in mind that the electronic density of Li is in a range where the hypothetical potential  $K_{xy}^{\text{byp}}$  introduced in our treatment of Na would be identical to  $K_{xy}^{\text{pz}}$ .

#### 6. Summary

In this work we calculated the volume dependence of the Knight shifts, K and the spin susceptibilities,  $\chi$ , of the light alkalines Na and Li for an extended volume range ( $V_{min}/V_0 = 0.65$  and 0.72 for Na and Li, respectively). A new formulation of the orbital part,  $K_{orb}$ , of the Knight shift has been derived and we evaluated all contributions to K in the frame of the SDA-RPA without introducing further approximations. These kind of investigations turned out to provide a sensitive test on the validity of the RPA and the accuracy of the expressions for the exchange-correlation energy functionals,  $E_{xc}$ , especially since the potentials  $K_{xc}$  are the second derivatives of  $E_{xc}$  with respect to the magnetization density.

In the case of Na the various potentials found in literature deviate appreciably from each other and lead to markedly different results for K and  $\chi$ . This feature is especially pronounced at low electronic densities. None of the potentials used could explain the minimum of K at medium pressures, observed by Bertani *et al* (1990). The differences among the potentials  $K_{xc}$  mainly near the surfaces of the Wigner-Seitz-spheres led us to the conclusion that they might be somewhat uncertain. Moreover, the potential  $K_{xc}^{PZ}$ which yields the best results for K is based on the Monte-Carlo calculations of Ceperley and Alder (1980). As these authors treat only the paramagnetic and the ferromagnetic state but not the case of partial magnetization,  $K_{xc}^{PZ}$  might be insufficiently accurate. By introducing the hypothetical potential,  $K_{xc}^{hyp}$ , we demonstrated that moderate changes of  $K_{xc}$  suffice to achieve agreement with experiment. Changes of this magnitude might well result from more accurate derivations of exchange correlation energy functional.

In the case of Li where the electronic density is higher and the various potentials,  $K_{xc}$ , show less scatter our results are in reasonable agreement with experiment.

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